

# Kinetic Modeling of Collisionless Magnetic Reconnection in Three Dimension

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# Outline

- Introduction
- Particle-in-Cell (PIC) model with adaptive mesh refinement (AMR)
- 3D kinetic modeling of reconnection  
*(Instabilities, turbulence, and anomalous transport at the x-line)*
- Summary

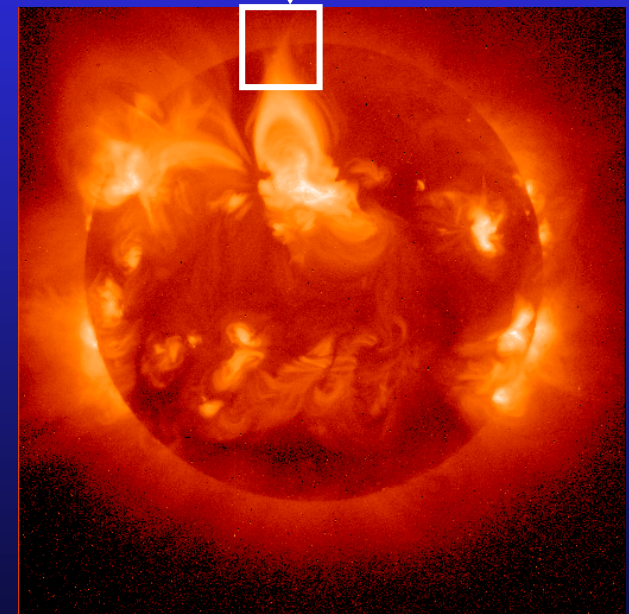
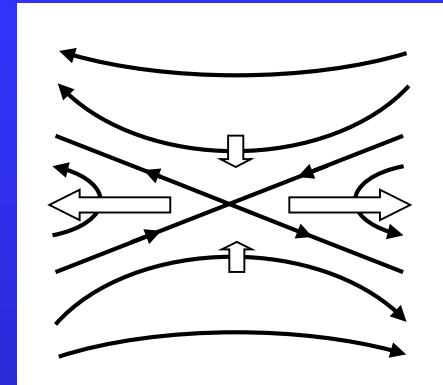
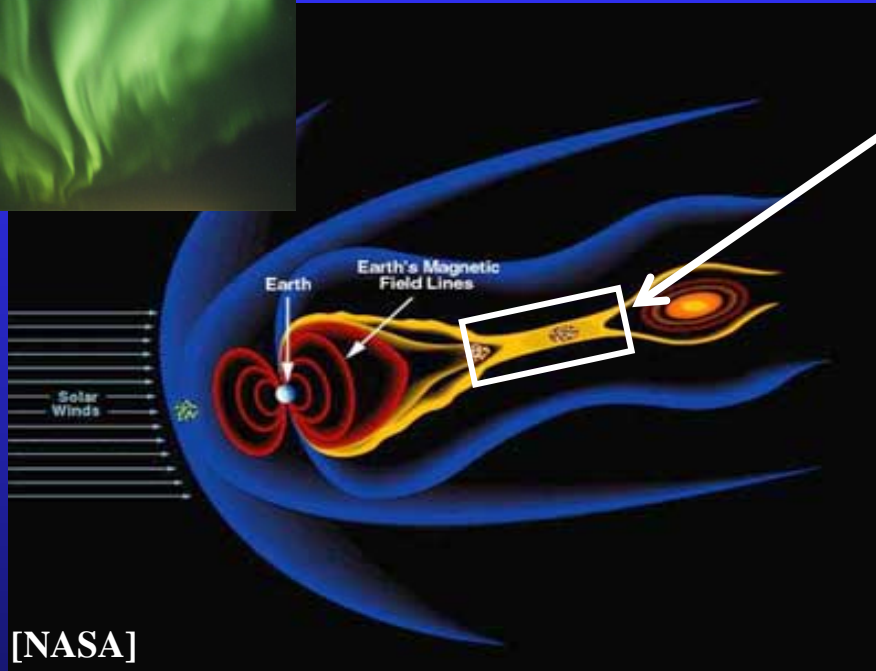
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# Magnetic Reconnection in Space

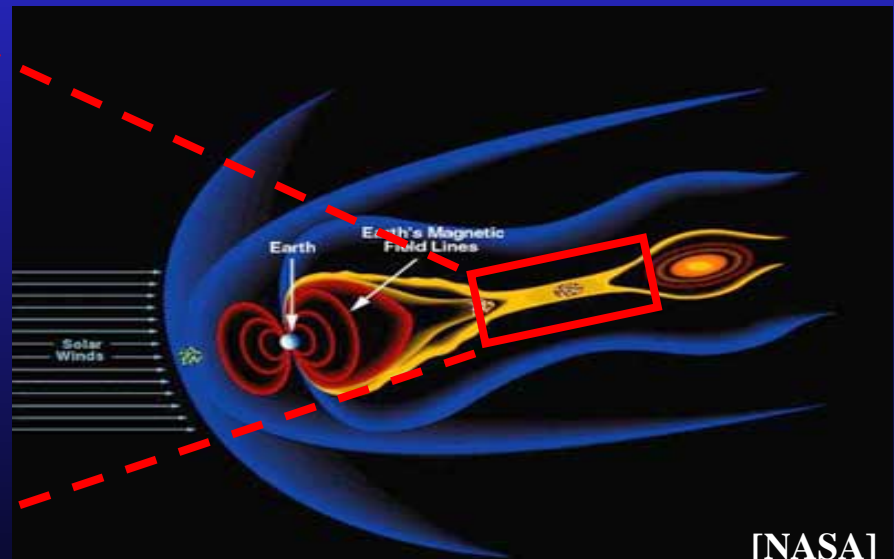
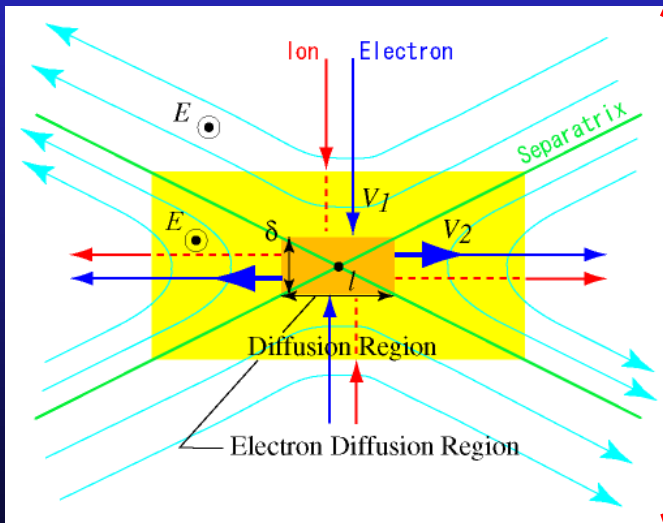
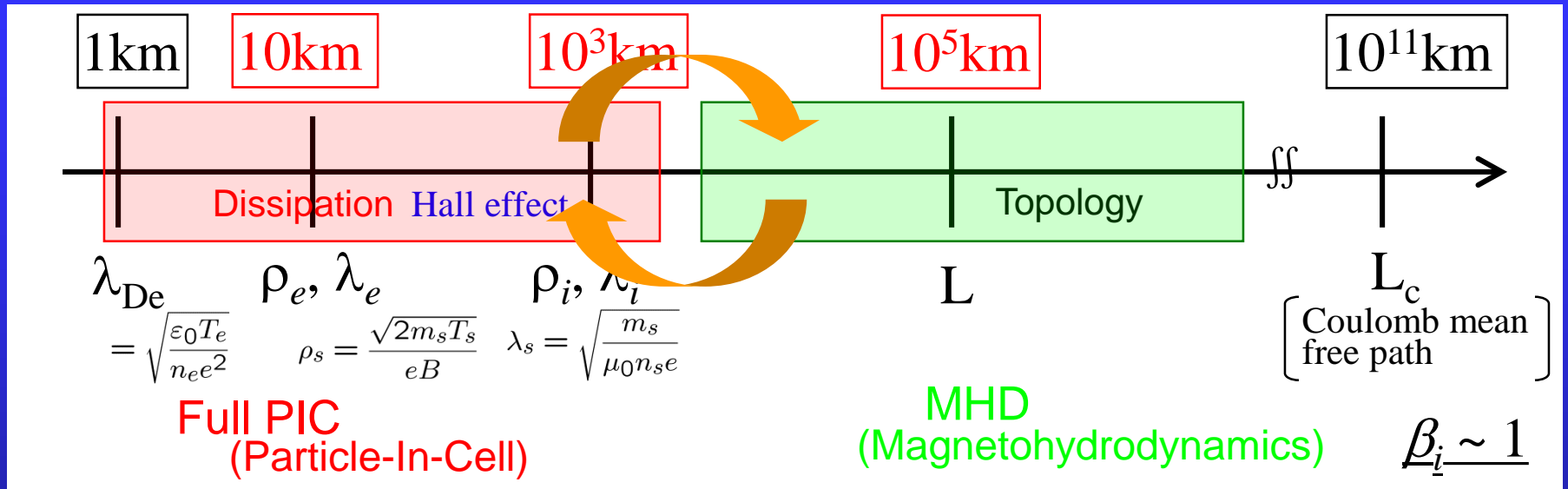


Auroral Substorms



Solar Flares

# Multi-Scale Nature of Reconnection



[NASA]

$$\frac{\partial B}{\partial t} = \eta \nabla^2 B / \mu_0$$

Numerical resistivity only

Nongyrotropic correction case

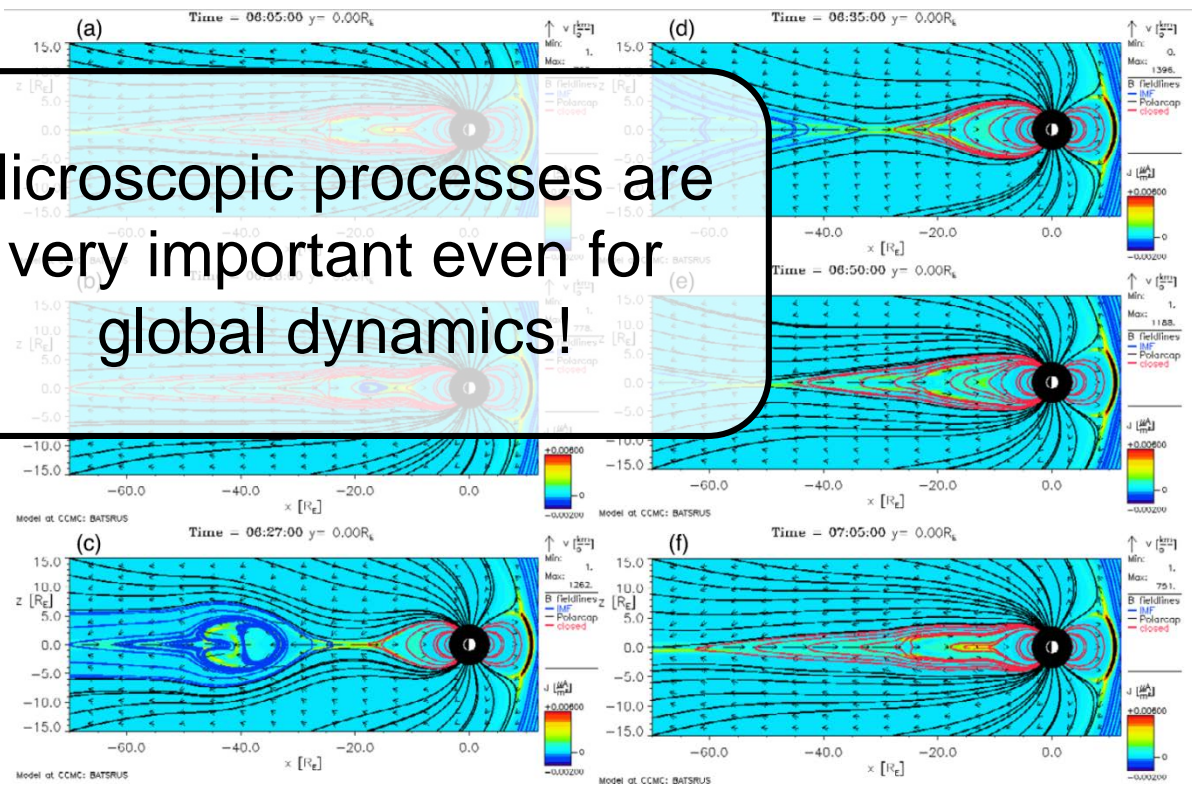
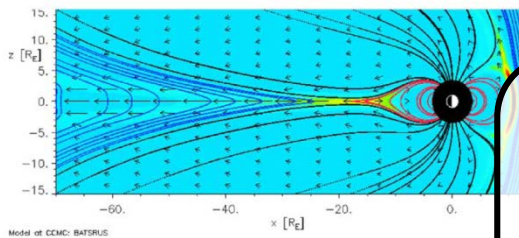
$$E^{ng} = \frac{1}{ne} \left( \frac{\partial P_{ixy}}{\partial x} + \frac{\partial P_{ixz}}{\partial z} \right) = \frac{m_i}{e} \sqrt{\frac{2P}{\rho}} \frac{\partial v_x}{\partial x}$$

Microscopic processes are very important even for global dynamics!

- Slow reconnection
- Quasi-steady configuration

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- Fast reconnection
- Quasi-periodic process



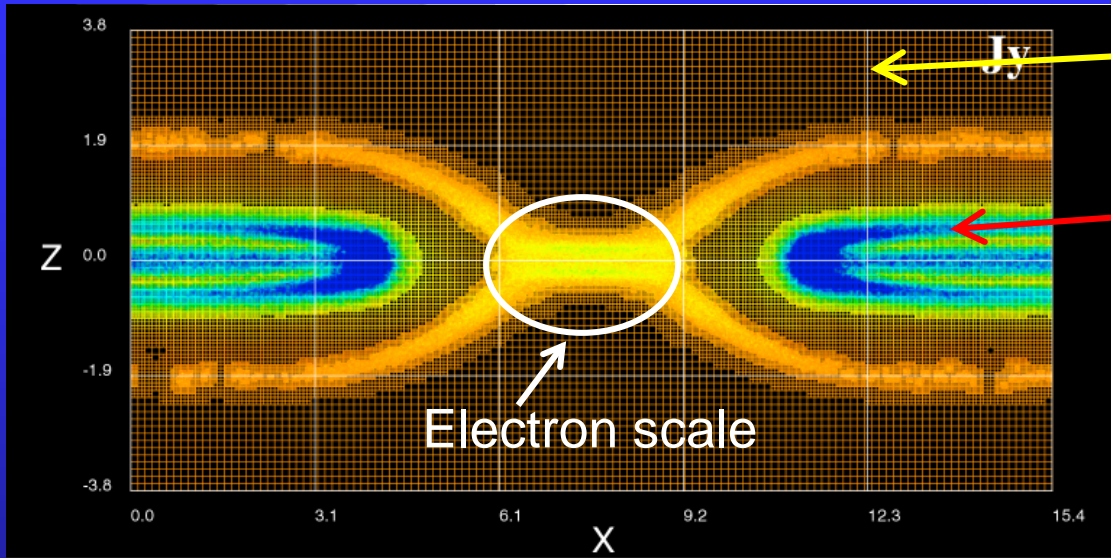
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# AMR-PIC Model [Fujimoto, JCP, 2011]

(Adaptive Mesh Refinement – Particle-in-Cell)

$$\Delta x \lesssim 3\lambda_{De} \propto n_e^{-1/2}$$

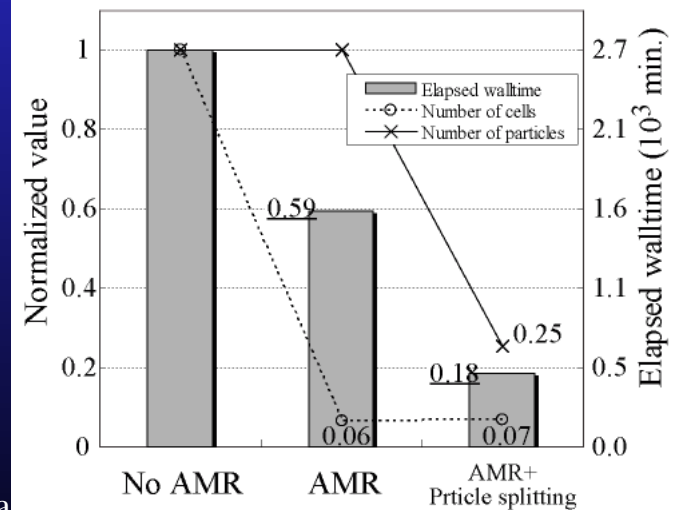
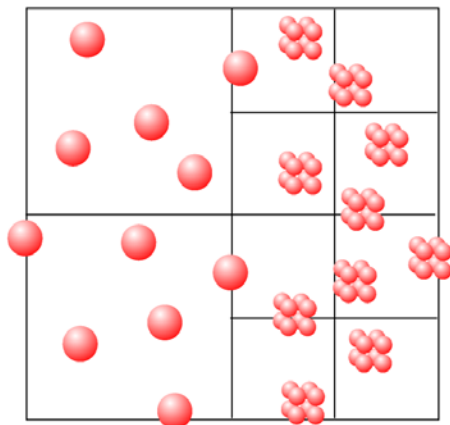


$$\lambda_{De,lobe} \sim 6 \times 10^3 \text{ m}$$

$$\lambda_{De,ps} \sim 3 \times 10^2 \text{ m}$$

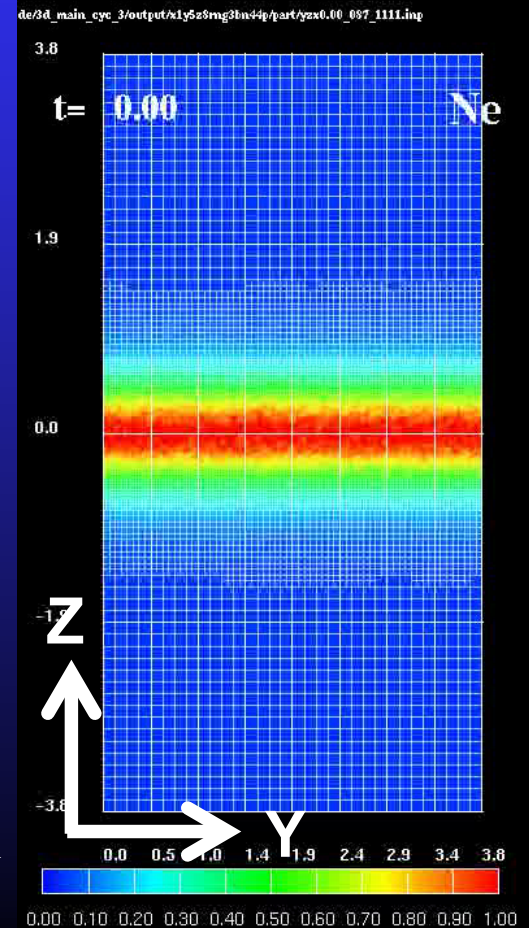
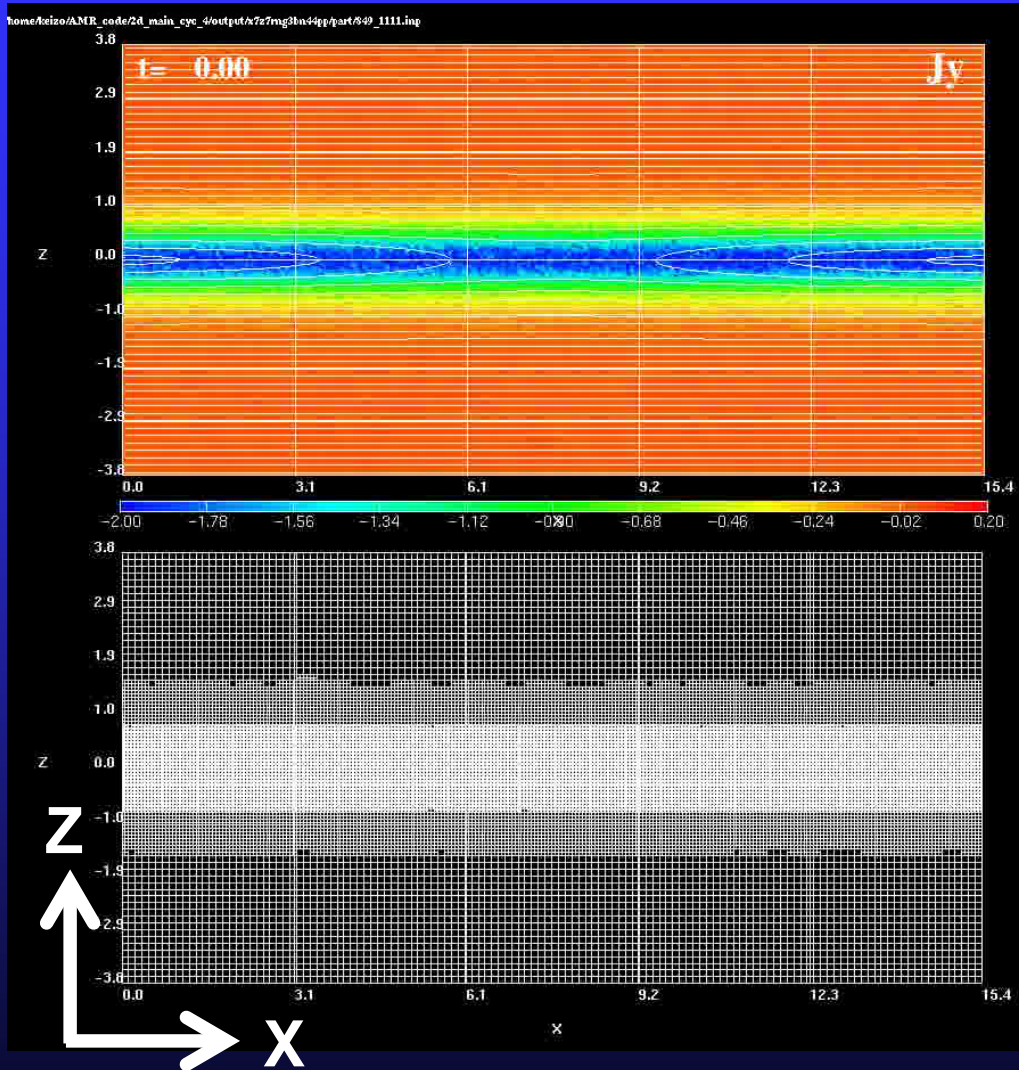
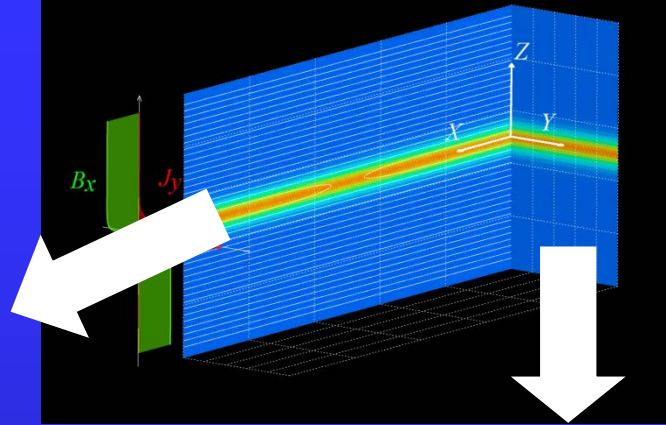
Removing unnecessary cells.

## Particle splitting-coalescence



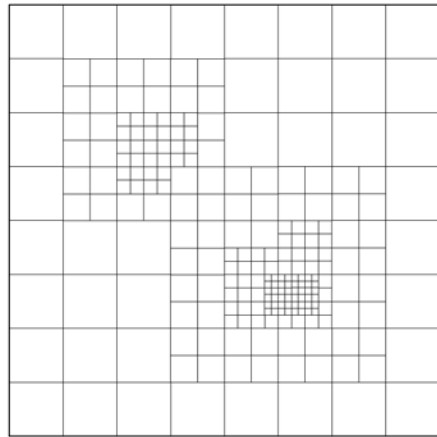


# AMR-PIC Simulations

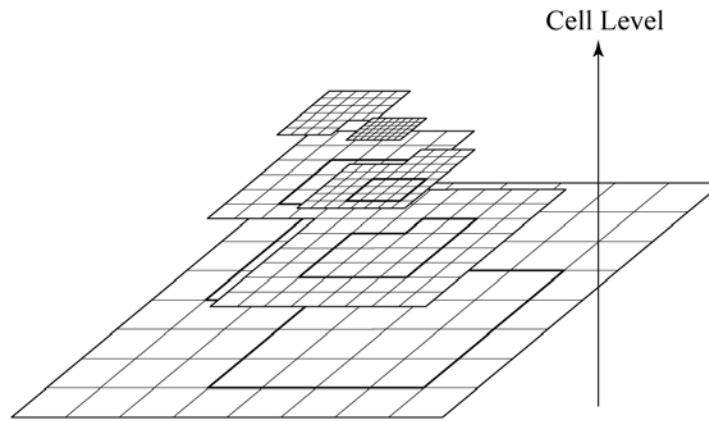


ASTRONUM2014@Long Beach, USA

# Data Structure



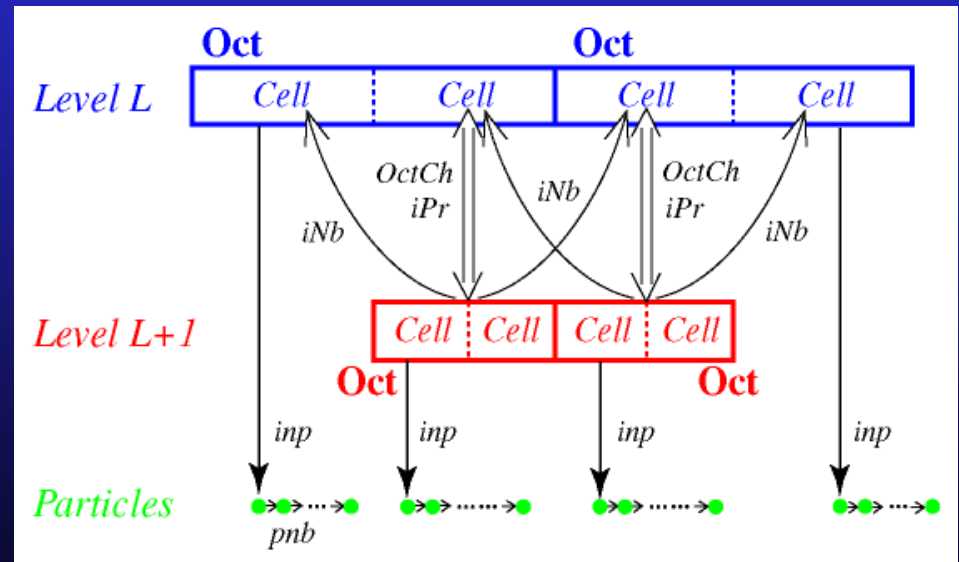
(a)



(b)

Similar to a fully threaded tree (FTT) structure (Khokhlov, 1998).

[Fujimoto & Machida, JCP, 2006]



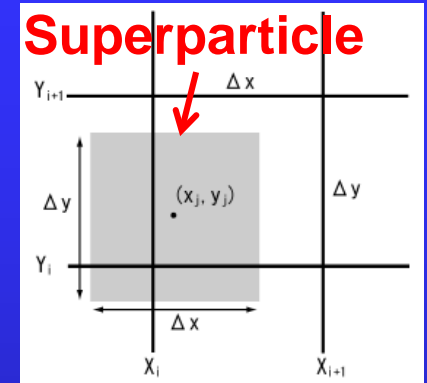
# Basic Equations

The same as usual PIC

$$\rho_{l,m,n} = \sum_s \sum_j q_{sj} S(\vec{x}_{sj} - \vec{X}_{l,m,n})$$

$$A(\vec{x}_{sj}) = \sum_l \sum_m \sum_n A_{l,m,n} S(\vec{x}_{sj} - \vec{X}_{l,m,n})$$

S: Shape function



## Superparticles (Buneman-Boris method)

$$\frac{\vec{v}_{sj}^{n+1/2} - \vec{v}_{sj}^{n-1/2}}{\Delta t} = \frac{q_{sj}}{m_{sj}} \left[ \vec{E}^n(\vec{x}_{sj}^n) + \frac{\vec{v}_{sj}^{n-1/2} + \vec{v}_{sj}^{n+1/2}}{2} \times \vec{B}^n(\vec{x}_{sj}^n) \right]$$

$$\frac{\vec{x}_{sj}^{n+1} - \vec{x}_{sj}^n}{\Delta t} = \vec{v}_{sj}^{n+1/2}$$

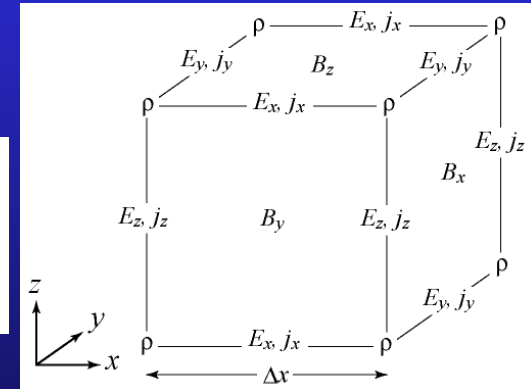
## EM Field (Yee-Buneman scheme)

$$\frac{\vec{B}^{n+1/2} - \vec{B}^{n-1/2}}{\Delta t} = -\nabla \times \vec{E}^n$$

$$\frac{\vec{E}^{n+1} - \vec{E}^n}{\Delta t} = c^2 \nabla \times \vec{B}^{n+1/2} - \frac{1}{\epsilon_0} \vec{j}^{n+1/2}$$

Charge Conservation Method

[Villasenor & Buneman, 1992]

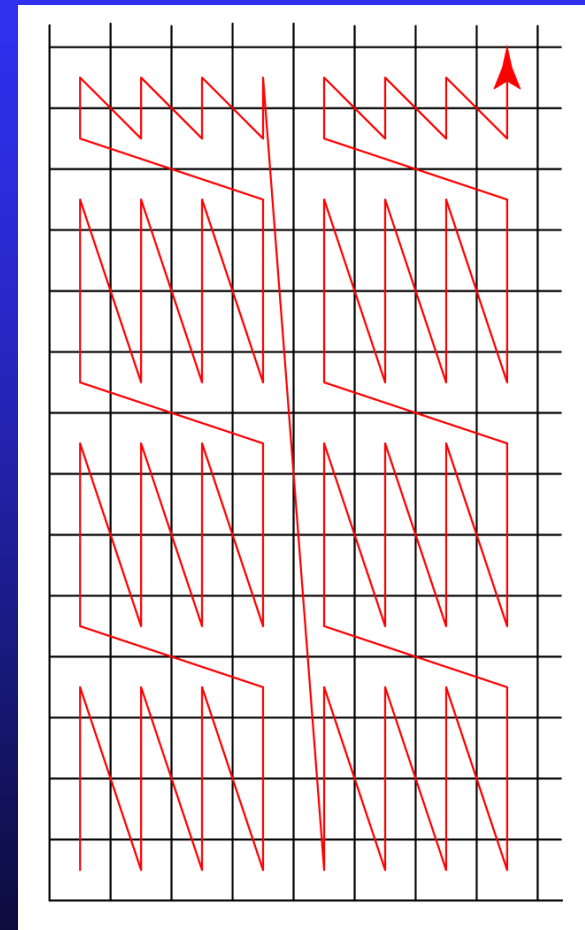
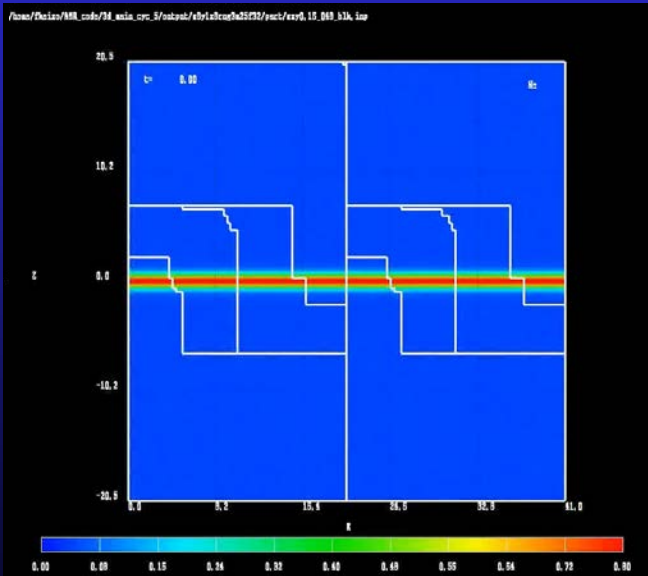


Local operations → Facilitate parallel computation

# Adaptive Block Technique [Fujimoto, JCP, 2011]

Base-level cells in the entire domain are sorted in an appropriate order:

- That is similar to Morton order,
- So that the block surface is as small as possible,
- Especially in the central current sheet, the surface must be small.



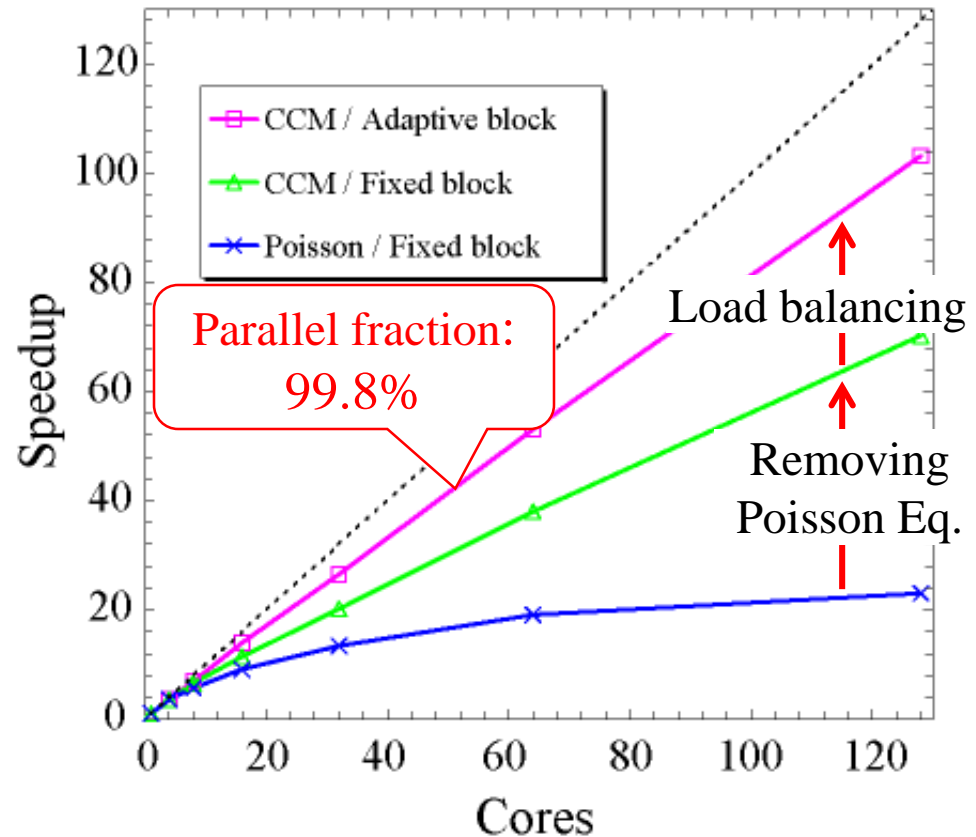
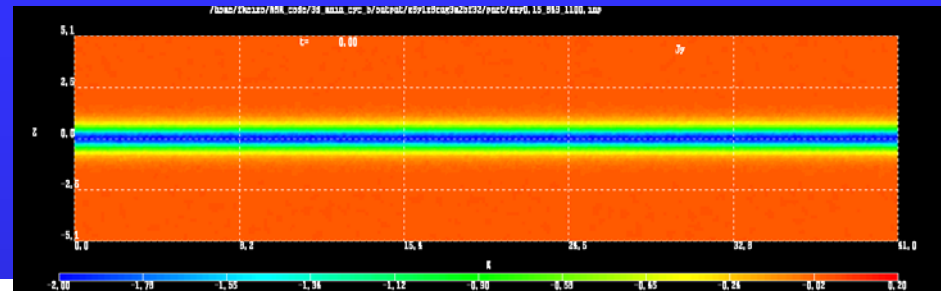
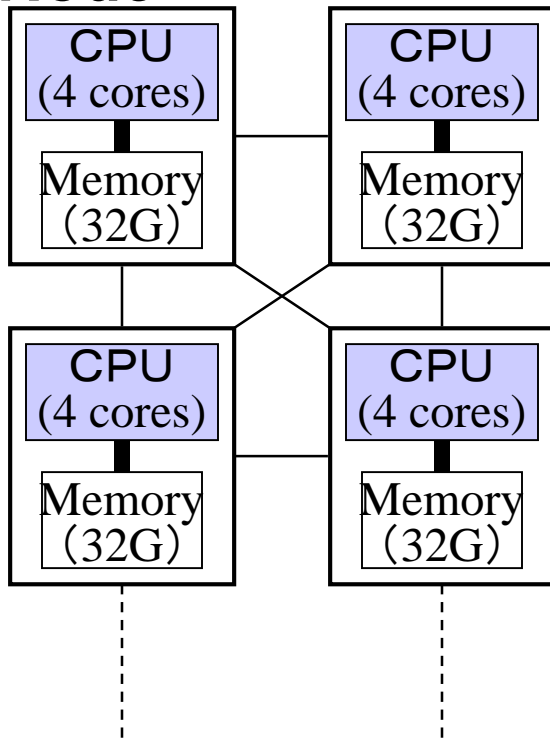
# Performance of the AMR-PIC Model

[Fujimoto, JCP, 2011]

Fujitsu FX1

@ Nagaya Univ.

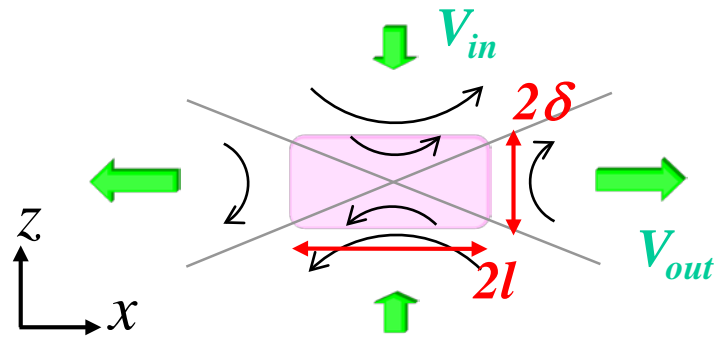
## Node



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*(Instabilities, turbulence, and anomalous transport at the x-line)*      Origin of resistivity...
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# Reconnection Rate and Resistivity $\eta$



Reconnection rate

$$E_R \equiv -\frac{\partial \Phi_m^*}{\partial t} \approx \frac{V_{in}}{V_{out}} \approx \frac{\delta}{l}$$

$$\Phi_m^* = \int_0^{L_z} B_x^* dz$$

[Vasyliunas, 1975]

Depending on width and length of the dissipation region.

$$\delta \approx \eta / \mu_0 V_{in} \quad \begin{array}{l} \text{2D case} \\ \rightarrow c / \omega_{pe} \end{array}$$

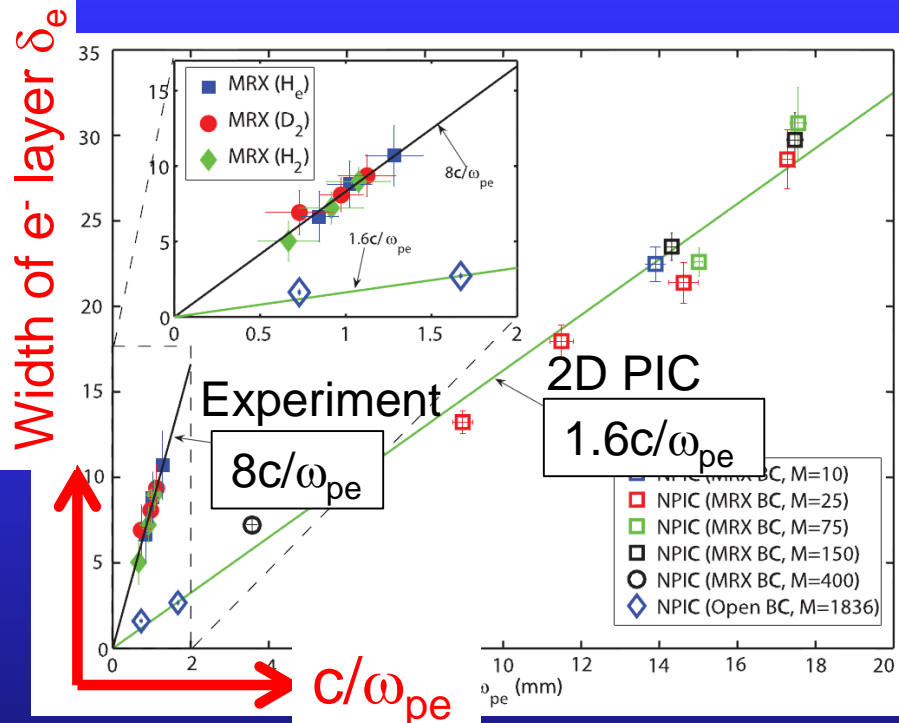
Inertia resistivity limit

3D, Multi scale, Nonlinear,  
Kinetic  $\Rightarrow$  Large-scale PIC

Inconsistent with  
Observations

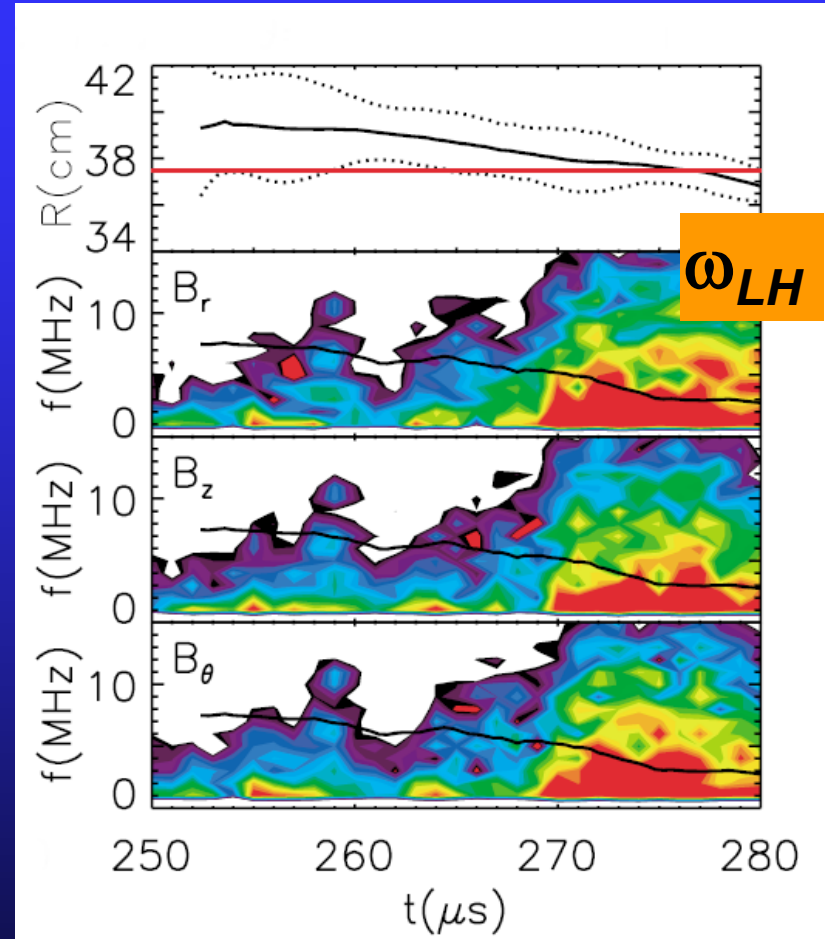
Anomalous dissipation?

# Laboratory Experiment



[Ji et al, GRL, 2008]

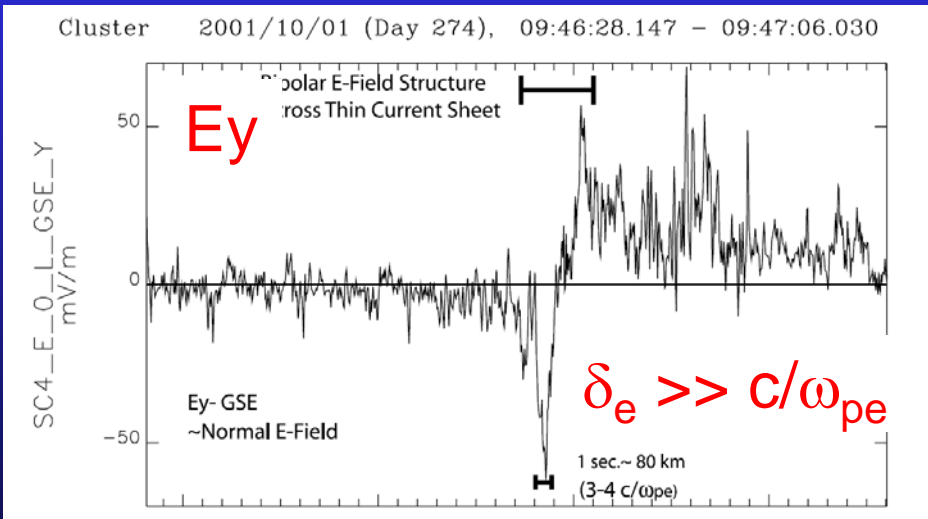
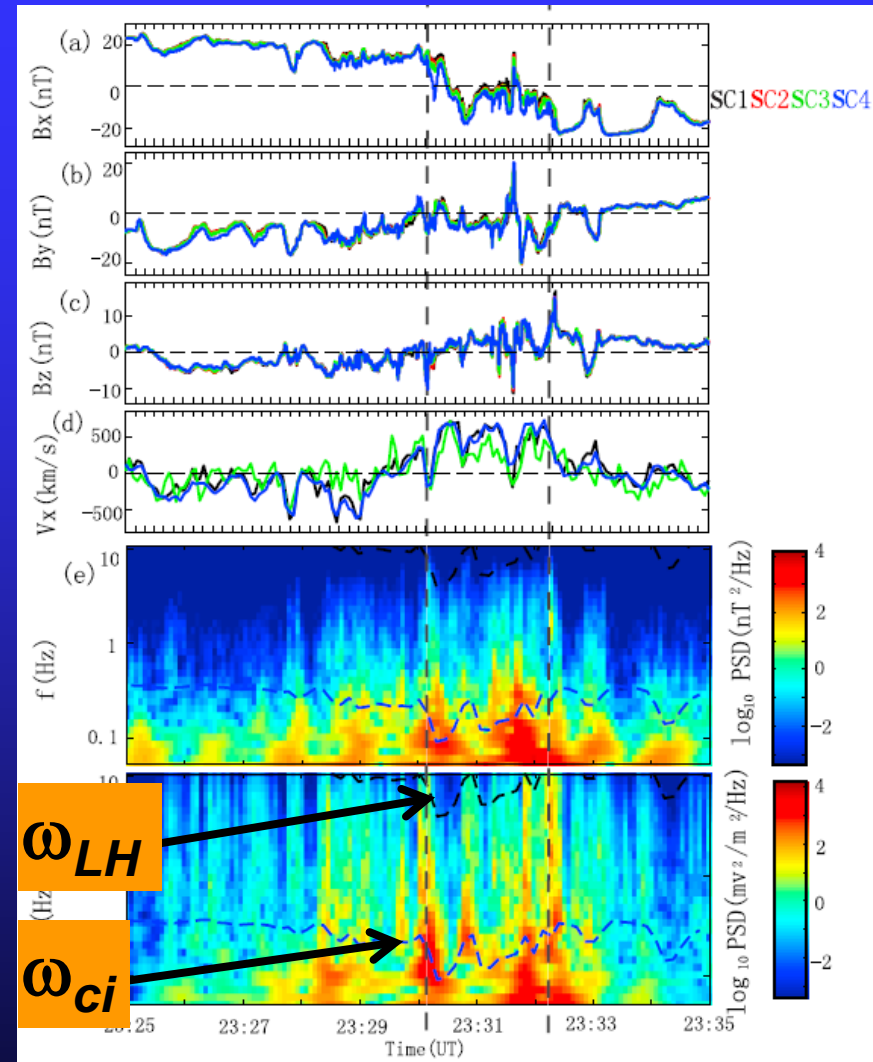
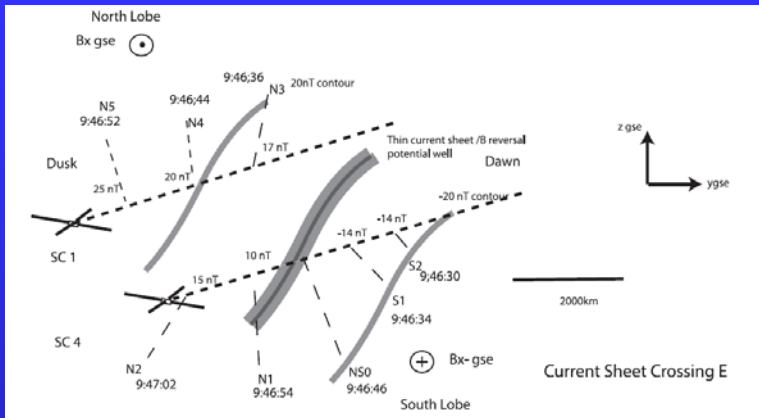
- $\delta_e \gg c/\omega_{pe}$
- Active  $\omega_{LH}$  waves



[Ji et al, PRL, 2004]



# Satellite Obs. in Earth's Magnetotail



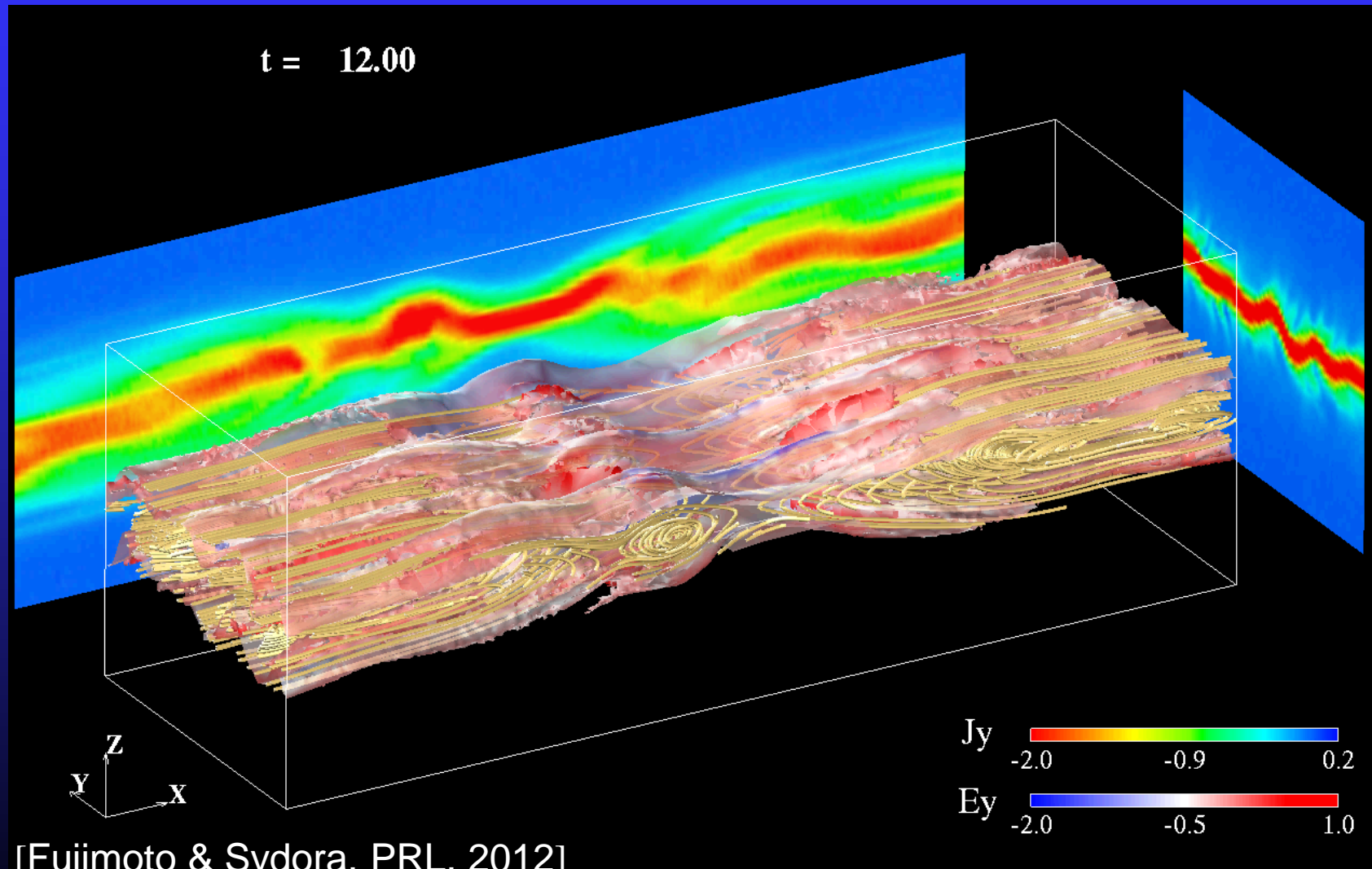
[Wygant et al, JGR, 2005]

[Zhou et al, JGR, 2009]

# 3D AMR-PIC of Anti-Parallel Reconnection

$m_i/m_e = 100$   $\sim 10^{11}$  particles  
 $\sim 6$  TB memory

Surface:  $|J|$ , Line: Field line  
Color on the surface:  $E_y$ , Cut plane:  $J_y$



# Generalized Ohm's Law

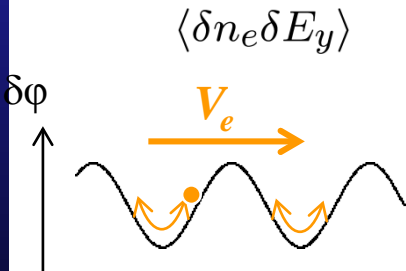
[Fujimoto & Sydora, PRL, 2012]

$$A = \langle A \rangle + \delta A \quad \left( \langle \cdot \rangle = \frac{1}{L_y} \int_0^{L_y} \cdot dy \right)$$

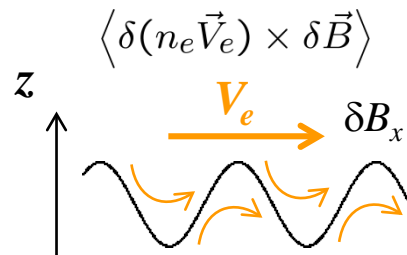
$$\begin{aligned} \langle -E_y \rangle &= \frac{1}{\langle n_e \rangle} \left( \langle n_e \vec{V}_e \rangle \times \langle \vec{B} \rangle \right)_y \\ &+ \frac{1}{e \langle n_e \rangle} \langle \nabla \cdot \vec{P}_e \rangle_y \\ &+ \frac{m_e}{e \langle n_e \rangle} \left\langle \frac{\partial V_{ey}}{\partial t} + \vec{V}_e \cdot \nabla V_{ey} \right\rangle \end{aligned}$$

$$\begin{aligned} &+ \frac{1}{\langle n_e \rangle} \langle \delta n_e \delta E_y \rangle \\ &+ \frac{1}{\langle n_e \rangle} \langle \delta(n_e \vec{V}_e) \times \delta \vec{B} \rangle_y \end{aligned}$$

Anomalous effects

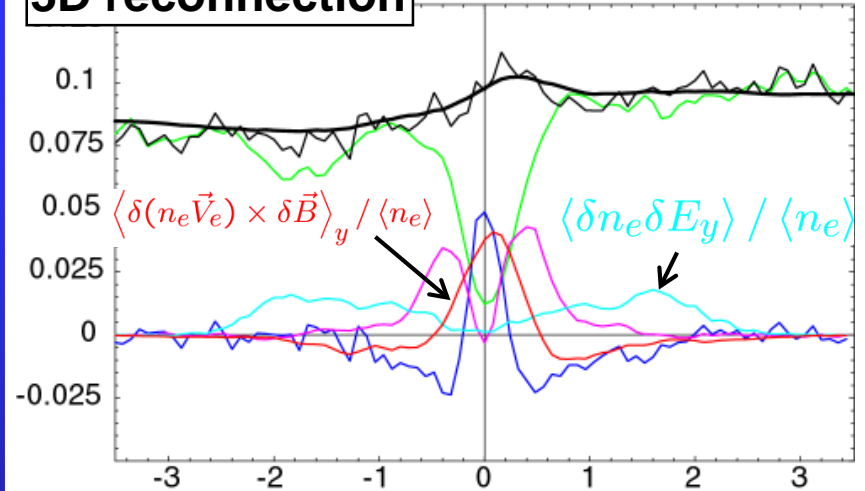


ES turb.

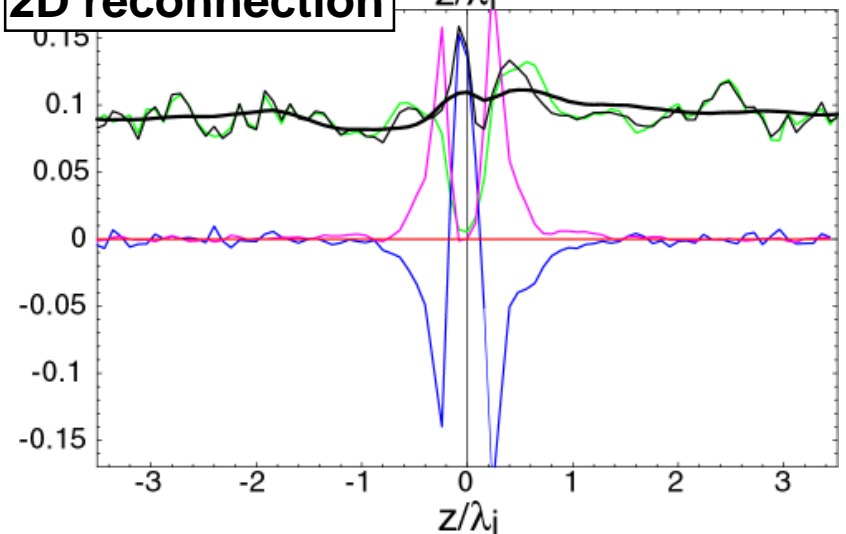


EM turb.

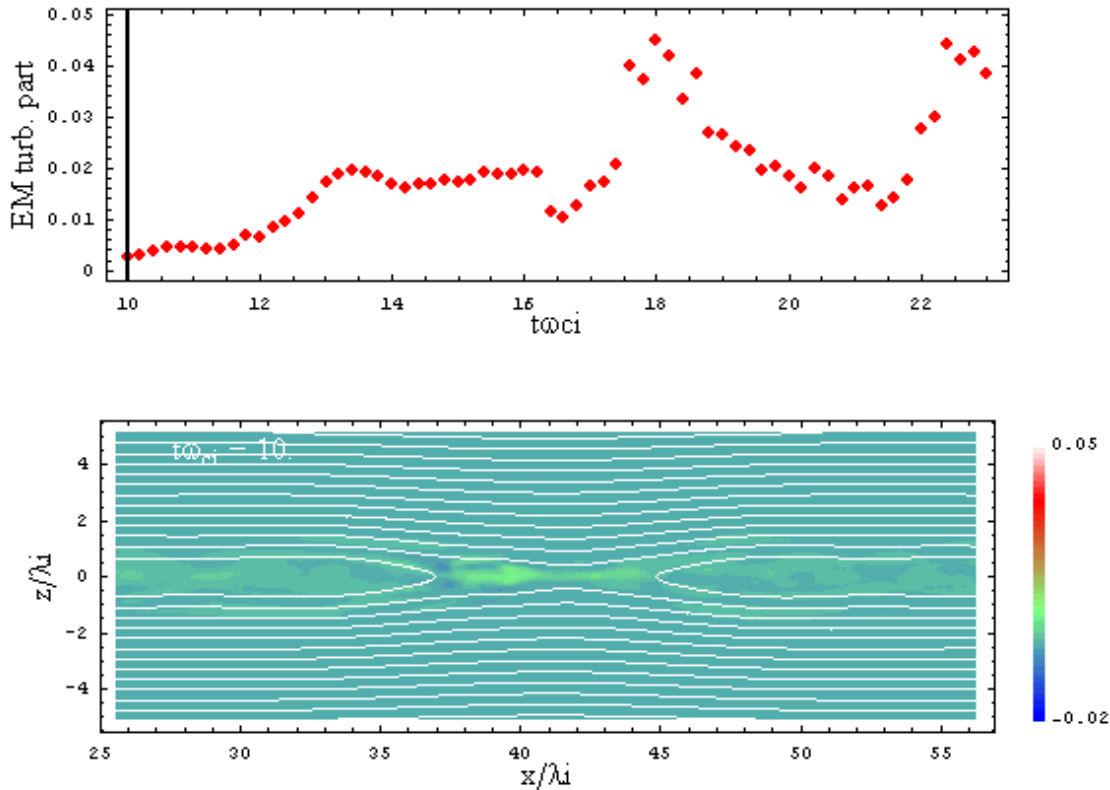
## 3D reconnection



## 2D reconnection



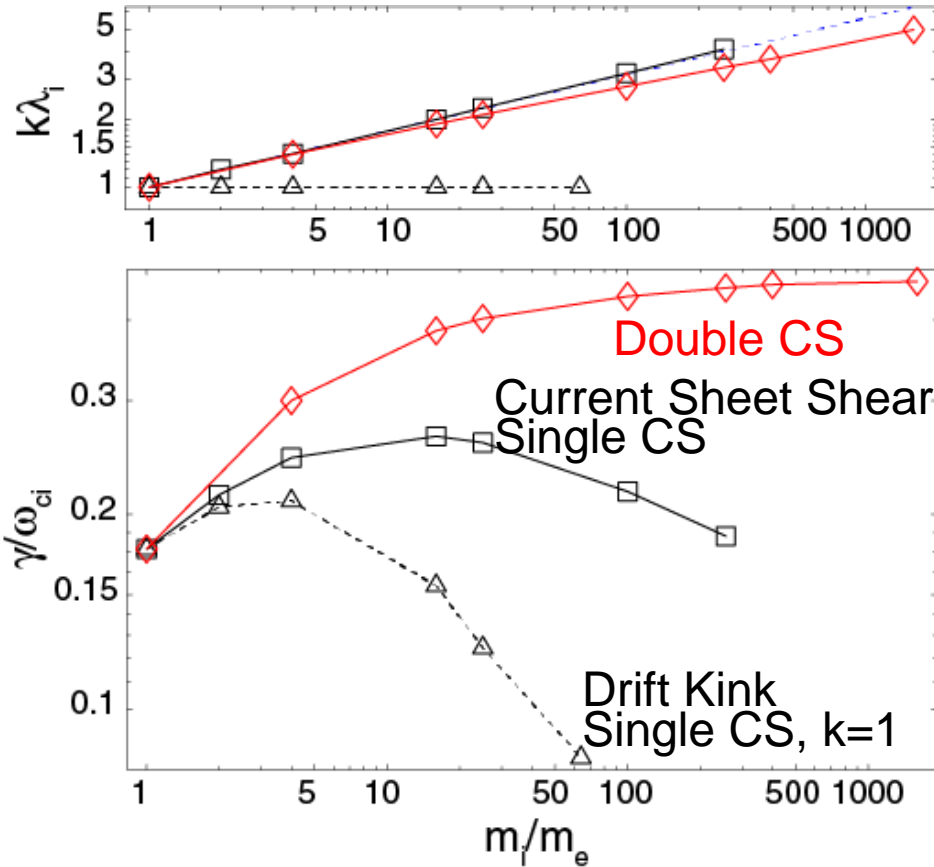
# Anomalous Transport at the X-line



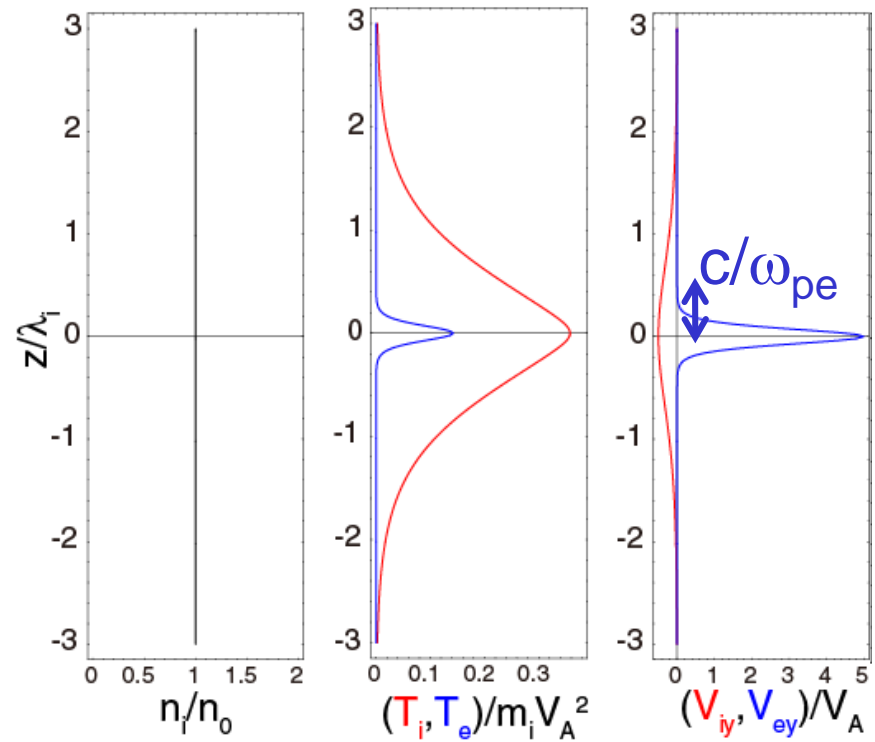
Turbulence is enhanced by plasmoid ejections.

# Linear Analyses of the Wave [Fujimoto & Sydora, in prep]

$$\omega = \omega_r + i\gamma$$

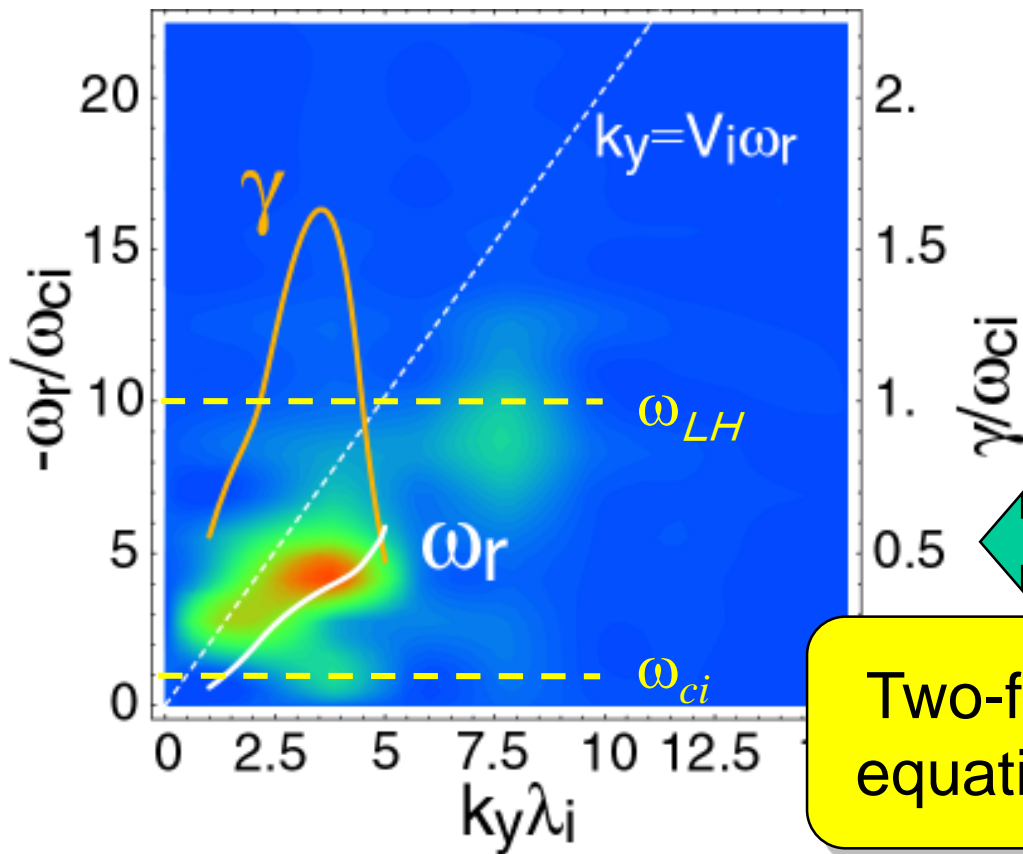


The mode is not the Drift Kink,  
not the LHDI,  
but the *Current Sheet Shear Inst.*



# Comparison with 3D Simulations

$$\omega = \omega_r + i\gamma$$



Two-fluid equations

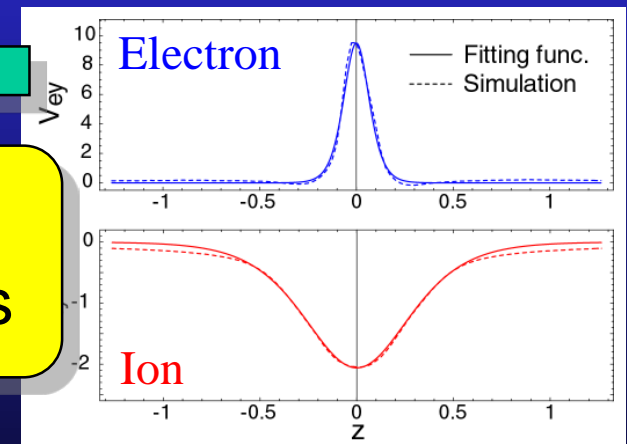
## Simulation results

$$\omega_{ci} < |\omega_r| < \omega_{LH}$$

$$V_{ph} \approx V_A$$

## Linear analyses

Profiles taken from simulation



# Summary

- Adaptive mesh refinement (AMR) has been implemented in the electromagnetic particle-in-cell (PIC) model to achieve large-scale simulations of magnetic reconnection.
- The AMR-PIC simulations indicate that the dissipation process of collisionless reconnection is three dimensional. The current sheet shear instability is dominant around the x-line. The turbulence is enhanced by the plasmoid ejections.